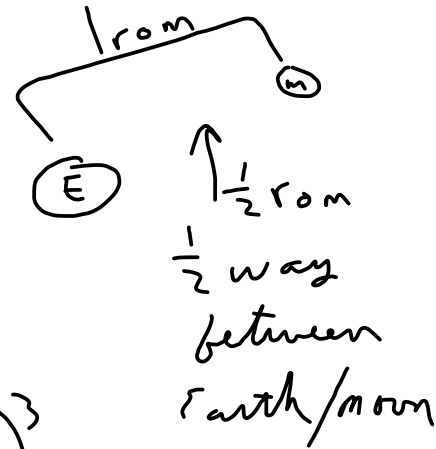


$$\frac{R_1^3}{T_1^2} = \frac{R_2^3}{T_2^2}$$

$$T_2 = ?$$



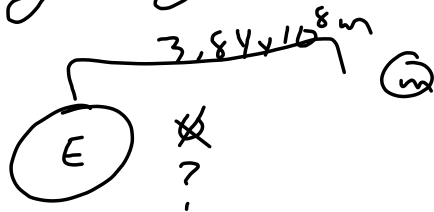
$$\frac{(1 \text{ rom})^3}{(27.3 \text{ da})^2} = \frac{(.5 \text{ rom})^3}{T_2^2}$$

$$T_2 = \sqrt{\frac{(.5 \text{ rom})^3 (27.3 \text{ da})^2}{(1 \text{ rom})^3}} = \underline{\underline{9.7 \text{ da}}}$$

$$\frac{\overset{\text{dist. to moon}}{(3.84 \times 10^8 \text{ m})^3}}{(2.36 \times 10^6 \text{ r})^2} = \frac{\overset{\substack{8.4 \times 10^5 \text{ r} \\ 1/2 \text{ dist to moon}}}{(1.92 \times 10^8 \text{ m})^3}}{T_2^2}$$

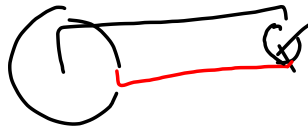
$$T_2 = 8.34 \times 10^5 \text{ r}$$

Where do you put a satellite to have a geosynchronous orbit.



$$\frac{r_m^3}{T_m^2} = \frac{r_s^3}{T_s^2} \quad \frac{(3.84 \times 10^8 \text{ m})^3}{(27.3 \text{ da})^2} = \frac{r_s^3}{(1 \text{ da})^2}$$

$$r_s = 4.23 \times 10^7 \text{ m}$$



from surface?

$$\begin{array}{r} 4.23 \times 10^7 \text{ m} \\ - .638 \times 10^7 \text{ m } r_e \\ \hline 3.592 \times 10^7 \text{ m} \end{array}$$

$$F = \frac{G m_1 m_2}{r^2}$$

Superior
← Inferior

$$m_2 a_c = F_c$$

$$\frac{m_2 v^2}{r} = F_c$$

$$\frac{m_2 4\pi^2 r}{T^2} = F_c$$

$$F = \frac{G m_1 m_2}{r^2}$$

$$m_2 a_c = F_c$$

~~$$m_2 a_c = \frac{G m_1 m_2}{r^2}$$~~



$$a_c = \frac{G m_1}{r^2}$$

What's the centripetal acceleration on the earth?

$$a_c = \frac{6.67 \times 10^{-11} \text{ N m}^2 / \text{kg}^2 (5.98 \times 10^{24} \text{ kg})}{(6.38 \times 10^6 \text{ m})^2}$$

$$a_c = 9.8 \text{ m/s}^2$$

...on the moon?

mass of moon

$$g_m = \frac{G (7.2 \times 10^{22} \text{ kg})}{(1.785 \times 10^6 \text{ m})^2}$$

$$g_m = 1.5 \text{ m/s}^2$$

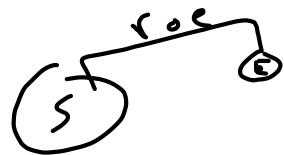
$$F = \frac{G m_1 m_2}{r^2}$$

$$\frac{m_2 v^2}{r} = F_c$$

$$\frac{\cancel{m_2} v^2}{\cancel{r}} = \frac{G m_1 \cancel{m_2}}{\cancel{r^2}}$$

$$v = \sqrt{\frac{G m_1}{r}}$$

What velocity does the earth circle the sun at?



$$v_e = \sqrt{\frac{G (1.99 \times 10^{30} \text{ kg})}{1.5 \times 10^{11} \text{ m}}}$$

mass of sun
↓
r_oe

$$v_e = 29,700 \text{ m/s}$$

$$2.97 \times 10^4 \text{ m/s}$$

$$F = \frac{G m_1 m_2}{r^2}$$

$$\frac{m_2 4\pi^2 r}{T^2} = F_c$$

$$\frac{\cancel{m_2} (4\pi^2) r}{T^2} = \frac{G m_1 \cancel{m_2}}{r^2}$$

Newton's variation of Kepler's 3rd Law

$$\frac{r^3}{T^2} = \frac{G m_1}{4\pi^2}$$

$$\frac{r^3}{T^2} = K$$